

Proxima Centauri

Proxima Centauri is a small, low-mass star located 4.244 light-years (1.301 pc) away from the Sun in the southern constellation of Centaurus. Its Latin name means the "nearest [star] of Centaurus". This object was discovered in 1915 by Robert Innes and is the nearest-known star to the Sun. With a quiescent apparent magnitude of 11.13, it is too faint to be seen with the naked eye. Proxima Centauri forms a third member of the Alpha Centauri system, being identified as component Alpha Centauri C, and is 2.18° to the southwest of the Alpha Centauri AB pair. Currently it has a physical separation of about 12,950 AU (1.94 trillion km) from AB and an orbital period of 550,000 years.

Proxima Centauri is a red dwarf star with a mass about an eighth of the Sun's mass (M_{\odot}), and average density about 33 times that of the Sun. Because of Proxima Centauri's proximity to Earth, its angular diameter can be measured directly as about one-seventh the diameter of the Sun. Although it has a very low average luminosity, Proxima is a flare star that undergoes random dramatic increases in brightness because of magnetic activity. The star's magnetic field is created by convection throughout the stellar body, and the resulting flare activity generates a total X-ray emission similar to that produced by the Sun. The mixing of the fuel at Proxima Centauri's core through convection and its relatively low energy-production rate mean that it will be a main-sequence star for another four trillion years.

In 2016, the European Southern Observatory announced the discovery of Proxima Centauri b, a planet orbiting the star at a distance of roughly 0.05 AU (7.5 million km) with an orbital period of approximately 11.2 Earth days. Its estimated mass is at least 1.3 times that of the Earth. The equilibrium temperature of Proxima b is estimated to be within the range of where water could exist as liquid on its surface, thus placing it within the habitable zone of Proxima Centauri, although because Proxima Centauri is a red dwarf and a flare star, whether it could support life is disputed.

Proxima Centauri



Proxima Centauri as seen by Hubble

Observation data

Epoch J2000.0 Equinox J2000.0 (ICRS)

Constellation	Centaurus
Pronunciation	/prəksɪmə sɛn'tɔ:rəɪ/ ^{[1][2]}
Right ascension	14 ^h 29 ^m 42.94853 ^s ^[3]
Declination	-62° 40' 46.1631" ^[3]
Apparent magnitude (V)	10.43 - 11.11 ^[4]
Characteristics	
Evolutionary stage	Main sequence red dwarf
Spectral type	M5.5Ve ^[5]
Apparent magnitude (U)	14.21 ^[6]
Apparent magnitude (B)	12.95 ^[6]
Apparent magnitude (V)	11.13 ^[6]
Apparent magnitude (R)	9.45 ^[6]
Apparent magnitude (I)	7.41 ^[6]
Apparent magnitude (J)	5.357 ± 0.023 ^[7]
Apparent magnitude (H)	4.835 ± 0.057 ^[7]
Apparent magnitude (K)	4.384 ± 0.033 ^[7]
U-B color index	1.26

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Observation

In 1915, the Scottish astronomer Robert Innes, Director of the Union Observatory in Johannesburg, South Africa, discovered a star that had the same proper motion as Alpha Centauri.^{[17][18][19][20]} He suggested that it be named *Proxima Centauri*^[21] (actually *Proxima Centaurus*).^[22] In 1917, at the Royal Observatory at the Cape of Good Hope, the Dutch astronomer Joan Voûte measured the star's trigonometric parallax at $0.755'' \pm 0.028''$ and determined that Proxima Centauri was approximately the same distance from the Sun as Alpha Centauri. It was also found to be the lowest-luminosity star known at the time.^[23] An equally accurate parallax determination of Proxima Centauri was made by American astronomer Harold L. Alden in 1928, who confirmed Innes's view that it is closer, with a parallax of $0.783'' \pm 0.005''$.^{[18][21]}



Stars closest to the Sun, including Proxima Centauri^[24]

In 1951, American astronomer Harlow Shapley announced that Proxima Centauri is a flare star. Examination of past photographic records showed that the star displayed a

B-V color index	1.82
V-R color index	1.68
R-I color index	2.04
J-H color index	0.522
J-K color index	0.973
Variable type	UV Ceti ("flare star") ^[4]
Astrometry	
Radial velocity (R_v)	-22.204 ± 0.032 ^[8] km/s
Proper motion (μ)	RA: -3781.306 ^[9] mas/yr Dec.: 769.766 ^[9] mas/yr
Parallax (π)	768.5 ± 0.2 ^[9] mas
Distance	4.244 ± 0.001 ly (1.3012 ± 0.0003 pc)
Absolute magnitude (M_V)	15.60 ^[10]
Orbit ^[8]	
Primary	Alpha Centauri AB
Companion	Proxima Centauri
Period (P)	$547\,000^{+6600}_{-4000}$ yr
Semi-major axis (a)	8700^{+700}_{-400} AU
Eccentricity (e)	$0.50^{+0.08}_{-0.09}$
Inclination (i)	$107.6^{+1.8^\circ}_{-2.0^\circ}$
Longitude of the node (Ω)	$126 \pm 5^\circ$
Periastron epoch (T)	$+283^{+59}_{-41}$
Argument of periastron (ω)	$72.3^{+8.7^\circ}_{-6.6^\circ}$ (secondary)
Details	
Mass	0.1221 ± 0.0022 ^[8] M _☉
Radius	0.1542 ± 0.0045 ^[8] R _☉
Luminosity (bolometric)	0.0017 ^[11] L _☉
Luminosity (visual, L_V)	$0.00005^{\text{[nb 1]}}$ L _☉
Surface gravity (log g)	5.20 ± 0.23 ^[12] cgs
Temperature	3042 ± 117 ^[12] K
Metallicity [Fe/H]	0.21 ^[13] dex
Rotation	82.6 ± 0.1 ^[14] days
Rotational velocity (v sin i)	< 0.1 ^[14] km/s
Age	4.85 ^[15] Gyr
Other designations	

measurable increase in magnitude on about 8% of the images, making it the most active flare star then known.^{[25][26]} The proximity of the star allows for detailed observation of its flare activity. In 1980, the [Einstein Observatory](#) produced a detailed X-ray energy curve of a stellar flare on Proxima Centauri. Further observations of flare activity were made with the [EXOSAT](#) and [ROSAT satellites](#), and the X-ray emissions of smaller, solar-like flares were observed by the Japanese [ASCA](#) satellite in 1995.^[27] Proxima Centauri has since been the subject of study by most X-ray observatories, including [XMM-Newton](#) and [Chandra](#).^[28]

In 2016, the [International Astronomical Union](#) organized a [Working Group on Star Names](#) (WGSN) to catalogue and standardize proper names for stars.^[29] The WGSN approved the name *Proxima Centauri* for this star on August 21, 2016 and it is now so included in the List of [IAU approved Star Names](#).^[30]

Because of Proxima Centauri's southern declination, it can only be viewed south of [latitude 27° N](#).^[nb 2] Red dwarfs such as Proxima Centauri are too faint to be seen with the naked eye. Even from Alpha Centauri A or B, Proxima would only be seen as a fifth magnitude star.^{[31][32]} It has an [apparent visual magnitude](#) of 11, so a [telescope with an aperture](#) of at least 8 cm (3.1 in) is needed to observe it, even under ideal viewing conditions—under clear, dark skies with Proxima Centauri well above the horizon.^[33]

In 2018, a superflare was observed from Proxima Centauri, the strongest flare ever seen. The optical brightness increased by a factor of 68 to approximately magnitude 6.8. It is estimated that similar flares occur around five times every year but are of such short duration, just a few minutes, that they have never been observed before.^[34]

Physical properties

Proxima Centauri is a [red dwarf](#), because it belongs to the [main sequence](#) on the [Hertzsprung–Russell diagram](#) and is of [spectral class M5.5](#). M5.5 means that it falls in the low-mass end of M-type stars.^[15] Its [absolute visual magnitude](#), or its visual magnitude as viewed from a distance of 10 parsecs (33 ly), is 15.5.^[35] Its total luminosity over all [wavelengths](#) is 0.17% that of the Sun,^[11] although when observed in the [wavelengths of visible light](#) the eye is most sensitive to, it is only 0.0056% as luminous as the Sun.^[36] More than 85% of its radiated power is at [infrared wavelengths](#).^[37] It has a regular activity cycle of [starspots](#).^[38]

In 2002, [optical interferometry](#) with the [Very Large Telescope](#) (VLTI) found that the [angular diameter](#) of Proxima Centauri is 1.02 ± 0.08 mas. Because its distance is known, the actual diameter of Proxima Centauri can be calculated to be about 1/7 that of the Sun, or 1.5 times that of [Jupiter](#). The star's mass, estimated from stellar theory, is 12.2% M_{\odot} , or 129 [Jupiter masses](#) (M_J).^[39] The mass has been calculated directly, although with less precision, from observations of [microlensing](#) events to be $0.150^{+0.062}_{-0.051} M_{\odot}$.^[40]

The mean [density](#) of main-sequence stars increase with decreasing mass,^[41] and Proxima Centauri is no exception: it has a mean density of 47.1×10^3 kg/m³ (47.1 g/cm³), compared with the Sun's mean density of 1.411×10^3 kg/m³ (1.411 g/cm³).^[nb 3]

A 1998 study of [photometric](#) variations indicates that Proxima Centauri rotates once every 83.5 days.^[42] A subsequent [time series](#) analysis of [chromospheric](#) indicators in 2002 suggests a longer rotation period of 116.6 ± 0.7 days.^[43] This was subsequently ruled out in favor of a rotation period of 82.6 ± 0.1 days.^[14]

Alpha Centauri C, CCDM J14396-

6050C, GCTP 3278.00, GJ 551,

HIP 70890, LFT 1110, LHS 49,

LPM 526, LTT 5721, NLTT 37460, V645

Centauri^[16]

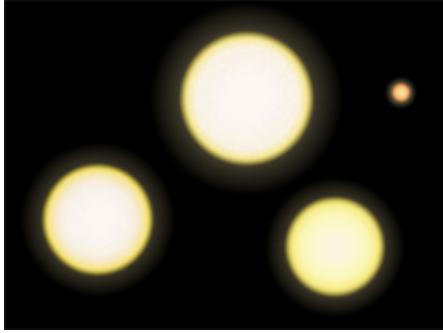
Database references

SIMBAD

data (<http://simbad.u-strasbg.fr/simbad/sim-id?ident=V645+Cen>)

ARICNS

data (<http://wwwadd.math.uni-heidelberg.de/datenbanken/aricns/cnspages/4c01140.htm>)



This illustration shows the comparative sizes of (from left to right) the Sun, Alpha Centauri A, Alpha Centauri B, and Proxima Centauri.



The two bright points are the Alpha Centauri system (left) and Beta Centauri (right). The faint red star in the centre of the red circle is Proxima Centauri.

4.85×10^9 years,^[15] since the activity level of a red dwarf is expected to steadily wane over billions of years as its stellar rotation rate decreases.^[50] The activity level also appears to vary with a period of roughly 442 days, which is shorter than the solar cycle of 11 years.^[51]

Proxima Centauri has a relatively weak stellar wind, no more than 20% of the mass loss rate of the solar wind. Because the star is much smaller than the Sun, the mass loss per unit surface area from Proxima Centauri may be eight times that from the solar surface.^[52]

A red dwarf with the mass of Proxima Centauri will remain on the main sequence for about four trillion years. As the proportion of helium increases because of hydrogen fusion, the star will become smaller and hotter, gradually transforming into a so-called "blue dwarf". Near the end of this period it will become significantly more luminous, reaching 2.5% of the Sun's luminosity (L_\odot) and warming up any orbiting bodies for a period of several billion years. When the hydrogen fuel is exhausted, Proxima Centauri will then evolve into a white dwarf (without passing through the red giant phase) and steadily lose any remaining heat energy.^[45]

Distance and motion

Based on a parallax of 768.5004 ± 0.2030 mas, published in 2018 in Gaia Data Release 2, Proxima Centauri is about 4.244 light-years (1.301 pc; 268,400 AU) from the Sun.^[9] Previously published parallaxes include: 768.13 ± 1.04 mas, in 2014 by the Research Consortium On Nearby Stars;^[53] 772.33 ± 2.42 mas, in the original Hipparcos Catalogue, in 1997;^[54] 771.64 ± 2.60 mas in the Hipparcos New Reduction, in 2007;^[3] and 768.77 ± 0.37 mas using the Hubble Space Telescope's Fine

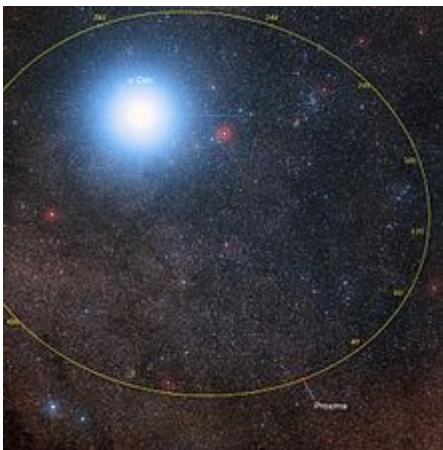
Because of its low mass, the interior of the star is completely convective,^[44] causing energy to be transferred to the exterior by the physical movement of plasma rather than through radiative processes. This convection means that the helium ash left over from the thermonuclear fusion of hydrogen does not accumulate at the core, but is instead circulated throughout the star. Unlike the Sun, which will only burn through about 10% of its total hydrogen supply before leaving the main sequence, Proxima Centauri will consume nearly all of its fuel before the fusion of hydrogen comes to an end after about 4 trillion years.^[45]

Convection is associated with the generation and persistence of a magnetic field. The magnetic energy from this field is released at the surface through stellar flares that briefly increase the overall luminosity of the star. These flares can grow as large as the star and reach temperatures measured as high as 27 million K^[28]—hot enough to radiate X-rays.^[46] Proxima Centauri's quiescent X-ray luminosity, approximately $(4\text{--}16) \times 10^{26}$ erg/s ($(4\text{--}16) \times 10^{19}$ W), is roughly equal to that of the much larger Sun. The peak X-ray luminosity of the largest flares can reach 10^{28} erg/s (10^{21} W).^[28]

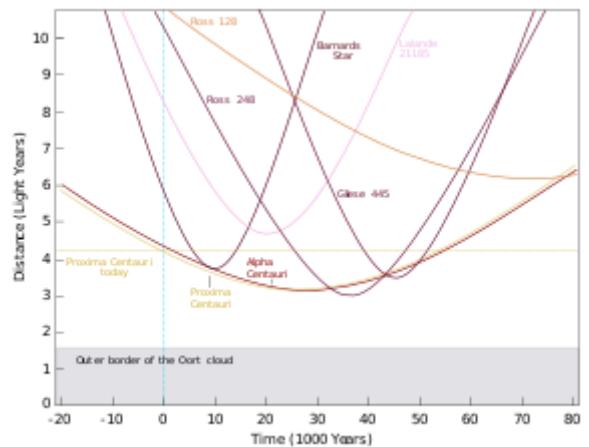
Proxima Centauri's chromosphere is active, and its spectrum displays a strong emission line of singly ionized magnesium at a wavelength of 280 nm.^[47] About 88% of the surface of Proxima Centauri may be active, a percentage that is much higher than that of the Sun even at the peak of the solar cycle. Even during quiescent periods with few or no flares, this activity increases the corona temperature of Proxima Centauri to 3.5 million K, compared to the 2 million K of the Sun's corona,^[48] and its total X-ray emission is comparable to the sun's.^[49] Proxima Centauri's overall activity level is considered low compared to other red dwarfs,^[49] which is consistent with the star's estimated age of

Guidance Sensors, in 1999.^[10] From Earth's vantage point, Proxima is separated from Alpha Centauri by 2.18 degrees,^[55] or four times the angular diameter of the full Moon.^[56] Proxima also has a relatively large proper motion—moving 3.85 arcseconds per year across the sky.^[57] It has a radial velocity toward the Sun of 22.2 km/s.^[8]

Among the known stars, Proxima Centauri has been the closest star to the Sun for about 32,000 years and will be so for about another 25,000 years, after which Alpha Centauri A and Alpha Centauri B will alternate approximately every 79.91 years as the closest star to the Sun. In 2001, J. García-Sánchez *et al.* predicted that Proxima will make its closest approach to the Sun in approximately 26,700 years, coming within 3.11 ly (0.95 pc).^[58] A 2010 study by V. V. Bobylev predicted a closest approach distance of 2.90 ly (0.89 pc) in about 27,400 years,^[59] followed by a 2014 study by C. A. L. Bailer-Jones predicting a perihelion approach of 3.07 ly (0.94 pc) in roughly 26,710 years.^[60] Proxima Centauri is orbiting through the Milky Way at a distance from the Galactic Centre that varies from 27 to 31 kly (8.3 to 9.5 kpc), with an orbital eccentricity of 0.07.^[61]



Orbital plot of Proxima Centauri as presently seen from Earth.^{[62][nb 4]}



Distances of the nearest stars from 20,000 years ago through 80,000 years in the future. Proxima Centauri is in yellow.

Ever since the discovery of Proxima, it has been suspected to be a true companion of the Alpha Centauri binary star system. Data from the Hipparcos satellite, combined with ground-based observations, were consistent with the hypothesis that the three stars are a bound system. For this reason, Proxima is sometimes referred to as Alpha Centauri C. Kervella *et al.* (2017) used high-precision radial velocity measurements to determine with a high degree of confidence that Proxima and Alpha Centauri are gravitationally bound.^[8] Proxima's orbital period around the Alpha Centauri AB barycenter is $547\,000^{+6600}_{-4000}$ years with an eccentricity of 0.5 ± 0.08 ; it approaches Alpha Centauri to 4300^{+1100}_{-900} AU at periastron and retreats to $13\,000^{+300}_{-100}$ AU at apastron.^[8] At present, Proxima is $12,947 \pm 260$ AU (1.94 ± 0.04 trillion km) from the Alpha Centauri AB barycenter, nearly to the farthest point in its orbit.^[8]

Such a triple system can form naturally through a low-mass star being dynamically captured by a more massive binary of $1.5\text{--}2 M_{\odot}$ within their embedded star cluster before the cluster disperses.^[63] However, more accurate measurements of the radial velocity are needed to confirm this hypothesis.^[64] If Proxima was bound to the Alpha Centauri system during its formation, the stars are likely to share the same elemental composition. The gravitational influence of Proxima might also have stirred up the Alpha Centauri protoplanetary disks. This would have increased the delivery of volatiles such as water to the dry inner regions, so possibly enriching any terrestrial planets in the system with this material.^[64] Alternatively, Proxima may have been captured at a later date during an encounter, resulting in a highly eccentric orbit that was then stabilized by the galactic tide and additional stellar encounters. Such a scenario may mean that Proxima's planetary companion has had a much lower chance for orbital disruption by Alpha Centauri.^[65]

Six single stars, two binary star systems, and a triple star share a common motion through space with Proxima Centauri and the Alpha Centauri system. The space velocities of these stars are all within 10 km/s of Alpha Centauri's peculiar motion. Thus, they may form a moving group of stars, which would indicate a common point of origin,^[66] such as in a star cluster.

Planetary system

The Proxima Centauri planetary system^{[67][68][69]}

Companion (in order from star)	Mass	Semimajor axis (AU)	Orbital period (days)	Eccentricity	Inclination	Radius
b	$\geq 1.27^{+0.19}_{-0.17} M_{\oplus}$	$0.0485^{+0.0041}_{-0.0051}$	11.186	<0.35	—	$0.8-1.5^{[70]} R_{\oplus}$
C (unconfirmed)	$\geq 6.0 \pm 1.9 M_{\oplus}$	1.48 ± 0.08	1894^{+92}_{-85}	—	—	—

The first indications of the exoplanet Proxima Centauri b were found in 2013 by Mikko Tuomi of the University of Hertfordshire from archival observation data.^{[71][72]} To confirm the possible discovery, the European Southern Observatory launched the Pale Red Dot^[nb 5] project in January 2016.^[73] On August 24, 2016, the team of 31 scientists from all around the world,^[74] led by Guillem Anglada-Escudé of Queen Mary University of London, confirmed the existence of Proxima Centauri b^[75] through a peer-reviewed article published in Nature.^{[67][76]} The measurements were performed using two spectrographs: HARPS on the ESO 3.6 m Telescope at La Silla Observatory and UVES on the 8 m Very Large Telescope at Paranal Observatory.^[67] Several attempts to detect a transit of this planet across the face of Proxima Centauri have been made. A transit-like signal appearing on September 8, 2016 was tentatively identified, using the Bright Star Survey Telescope at the Zhongshan Station in Antarctica.^[77]

Proxima Centauri b, or Alpha Centauri Cb, is a planet orbiting the star at a distance of roughly 0.05 AU (7.5 million km) with an orbital period of approximately 11.2 Earth days. Its estimated mass is at least 1.3 times that of the Earth. Moreover, the equilibrium temperature of Proxima b is estimated to be within the range where water could exist as liquid on its surface; thus, placing it within the habitable zone of Proxima Centauri.^{[67][79][80]}

RV-derived upper mass limits of potential companions^[78]

Orbital period (days)	Separation (AU)	Maximum mass ^[nb 6] (M_{\oplus})
3.6–13.8	0.022–0.054	2–3
<100	<0.21	8.5
<1000	<1	16

Prior to this discovery, multiple measurements of the star's radial velocity constrained the maximum mass that a detectable companion to Proxima Centauri could possess.^{[10][81]} The activity level of the star adds noise to the radial velocity measurements, complicating detection of a companion using this method.^[82] In 1998, an examination of Proxima Centauri using the Faint Object Spectrograph on board the Hubble Space Telescope appeared to show evidence of a companion orbiting at a distance of about 0.5 AU.^[83] A subsequent search using the Wide Field Planetary Camera 2 failed to locate any companions.^[84] Astrometric measurements at the Cerro Tololo Inter-American Observatory appear to rule out a Jupiter-sized planet with an orbital period of 2–12 years.^[85] A second signal in the range of 60 to 500 days was also detected, but its nature is still unclear due to stellar activity.^[67]

A candidate second planet orbiting Proxima Centauri was reported by Italian astrophysicist Mario Damasso and his colleagues in April 2019.^{[86][87]} Damasso's team had noticed minor movements of Proxima Centauri in the radial velocity data from the ESO's HARPS instrument, indicating a possible additional planet orbiting Proxima Centauri.^[86] Dubbed Proxima c, the exoplanet is estimated to have a minimum mass of six times that of Earth.^[86] It is expected to orbit Proxima Centauri at a distance of 1.5 AU, with an orbital period of roughly 1900 days, or about 5.2 years.^[87] Due to its large distance from Proxima Centauri, the exoplanet is unlikely to be habitable, with a low equilibrium temperature of around 39 K.^[86] Additional observations and measurements from the HARPS instrument and the European Space Agency's Gaia spacecraft are needed to verify the existence of the possible exoplanet.^[86] Del Sordo of Damasso's team states that Proxima c could provide opportunities for further observations of the Proxima Centauri planetary system, especially by direct imaging.^{[86][87]}

Proxima Centauri, along with Alpha Centauri A and B, was among the "Tier 1" target stars for NASA's now-canceled Space Interferometry Mission (SIM), which would theoretically have been able to detect planets as small as three Earth masses (M_{\oplus}) within two AU of a "Tier 1" target star.^[88]

In 2017, a team of astronomers using the Atacama Large Millimeter/submillimeter Array reported detecting a belt of dust orbiting Proxima Centauri at a range of 1–4 AU from the star. This dust has a temperature of around 40 K and has a total estimated mass of 1% of the planet Earth. They also tentatively detected two additional features: a cold belt with a temperature of 10 K orbiting around 30 AU and a compact emission source about 1.2 arcseconds from the star.^[89] However, upon further analysis, these emissions were determined to be the result of a large flare emitted by the star in March, 2017. The presence of dust is not needed to model the observations.^{[90][91]}

Habitability

Prior to the discovery of Proxima Centauri b, the TV documentary *Alien Worlds* hypothesized that a life-sustaining planet could exist in orbit around Proxima Centauri or other red dwarfs. Such a planet would lie within the habitable zone of Proxima Centauri, about 0.023–0.054 AU (3.4–8.1 million km) from the star, and would have an orbital period of 3.6–14 days.^[93] A planet orbiting within this zone may experience tidal locking to the star. If the orbital eccentricity of this hypothetical planet is low, Proxima Centauri would move little in the planet's sky, and most of the surface would experience either day or night perpetually. The presence of an atmosphere could serve to redistribute the energy from the star-lit side to the far side of the planet.^[94]



Pale Red Dot is an international search for an Earth-like exoplanet around the closest star Proxima Centauri.^[92]

Proxima Centauri's flare outbursts could erode the atmosphere of any planet in its habitable zone, but the documentary's scientists thought that this obstacle could be overcome. Gibor Basri of the University of California, Berkeley, mentioned that "no one [has] found any showstoppers to habitability". For example, one concern was that the torrents of charged particles from the star's flares could strip the atmosphere off any nearby planet. If the planet had a strong magnetic field, the field would deflect the particles from the atmosphere; even the slow rotation of a tidally locked planet that spins once for every time it orbits its star would be enough to generate a magnetic field, as long as part of the planet's interior remained molten.^[95]

Other scientists, especially proponents of the rare-Earth hypothesis,^[96] disagree that red dwarfs can sustain life. Any exoplanet in this star's habitable zone would likely be tidally locked, resulting in a relatively weak planetary magnetic moment, leading to strong atmospheric erosion by coronal mass ejections from Proxima Centauri.^[97]

Future exploration

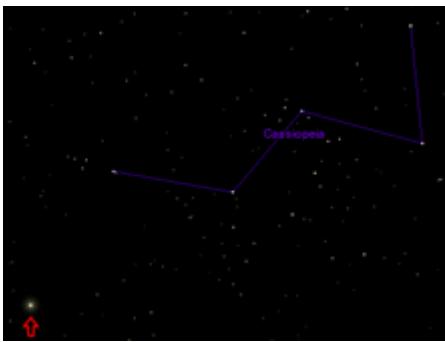
Because of the star's proximity to Earth, Proxima Centauri has been proposed as a flyby destination for interstellar travel.^[98] Proxima currently moves toward Earth at a rate of 22.2 km/s.^[8] After 26,700 years, when it will come within 3.11 light-years, it will begin to move farther away.^[58]

If non-nuclear, conventional propulsion technologies are used, the flight of a spacecraft to a planet orbiting Proxima Centauri would probably require thousands of years.^[99] For example, Voyager 1, which is now travelling 17 km/s (38,000 mph)^[100] relative to the Sun, would reach Proxima in 73,775 years, were the spacecraft travelling in the direction of that star. A slow-moving probe would have only several tens of thousands of years to catch Proxima Centauri near its closest approach, and could end up watching it recede into the distance.^[101]

Nuclear pulse propulsion might enable such interstellar travel with a trip timescale of a century, inspiring several studies such as Project Orion, Project Daedalus, and Project Longshot.^[101]

Project Breakthrough Starshot aims to reach the Alpha Centauri system within the first half of the 21st century, with microprobes travelling at 20% of the speed of light propelled by around 100 gigawatts of Earth-based lasers.^[102] The probes would perform a fly-by of Proxima Centauri to take photos and collect data of its planet's atmospheric composition. It would take 4.22 years for the information collected to be sent back to Earth.^[103]

From Proxima Centauri, the Sun would appear as a bright 0.4-magnitude star in the constellation Cassiopeia, similar to that of Achernar from Earth.^[nb 7]



The Sun as seen from the Alpha Centauri system, using Celestia.

See also

- Proxima Centauri in fiction

Notes

1. From knowing the absolute visual magnitude of Proxima Centauri, $M_{V*}=15.6$, and the absolute visual magnitude of the Sun, $M_{V\odot}=4.83$, the visual luminosity of Proxima Centauri can therefore be calculated: $\frac{L_{V*}}{L_{V\odot}}=10^{0.4(M_{V\odot}-M_{V*})}=4.92\times10^{-5}$
2. For a star south of the zenith, the angle to the zenith is equal to the Latitude minus the Declination. The star is hidden from sight when the zenith angle is 90° or more, i.e. below the horizon. Thus, for Proxima Centauri:

$$\text{Highest latitude} = 90^\circ + -62.68^\circ = 27.32^\circ.$$

See: Campbell, William Wallace (1899). *The elements of practical astronomy* (<https://books.google.com/?id=v2tEAAAAIAAJ>). London: Macmillan. pp. 109–110. Retrieved August 12, 2008.

3. The density (ρ) is given by the mass divided by the volume. Relative to the Sun, therefore, the density is:

$$\begin{aligned}\rho &= \frac{M}{M_\odot} \cdot \left(\frac{R}{R_\odot}\right)^{-3} \cdot \rho_\odot \\ &= 0.122 \cdot 0.154^{-3} \cdot (1.41 \times 10^3 \text{ kg/m}^3) \\ &= 33.4 \cdot (1.41 \times 10^3 \text{ kg/m}^3) \\ &= 4.71 \times 10^4 \text{ kg/m}^3\end{aligned}$$

where ρ_\odot is the average solar density. See:

- Munsell, Kirk; Smith, Harman; Davis, Phil; Harvey, Samantha (June 11, 2008). "Sun: facts & figures" (<https://web.archive.org/web/20080102034758/http://solarsystem.nasa.gov/planets/profile.cfm?Object=Sun&Display=Facts&System=Metric>). *Solar system exploration*. NASA. Archived from the original (<http://solarsystem.nasa.gov/planets/profile.cfm?Object=Sun&Display=Facts&System=Metric>) on January 2, 2008. Retrieved July 12, 2008.
 - Bergman, Marcel W.; Clark, T. Alan; Wilson, William J. F. (2007). *Observing projects using Starry Night Enthusiast* (8th ed.). Macmillan. pp. 220–221. ISBN 978-1-4292-0074-5.
4. Note that by the time Proxima gets to the 40,000-year mark, the entire Alpha Centauri system will have moved to another part of the sky, so the perspective and background will be different.
 5. Pale Red Dot is a reference to Pale Blue Dot, a distant photo of Earth taken by Voyager 1.

6. This is actually an upper limit on the quantity $m \sin i$, where i is the angle between the orbit normal and the line of sight, in a circular orbit. If the planetary orbits are close to face-on as observed from Earth, or in an eccentric orbit, more massive planets could have evaded detection by the radial velocity method.
7. The coordinates of the Sun would be diametrically opposite Proxima, at $\alpha = 02^{\text{h}} 29^{\text{m}} 42.9487^{\text{s}}$, $\delta = +62^{\circ} 40' 46.141''$. The absolute magnitude M_V of the Sun is 4.83, so at a parallax π of 0.77199 the apparent magnitude m is given by $4.83 - 5(\log_{10}(0.77199) + 1) = 0.40$. See: Tayler, Roger John (1994). *The Stars: Their Structure and Evolution*. Cambridge University Press. p. 16. [ISBN 978-0-521-45885-6](#).

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